

Physics 12c, Problem Set 8 Solutions

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[1] Hyperscaling

For $\lambda = 0$, the scaling hypothesis becomes

$$G(\Omega^p \epsilon) = \Omega G(\epsilon);$$

hence choosing $\Omega^p \epsilon = \text{constant}$, or $\Omega \sim |\epsilon|^{-1/p}$, we find

$$G(\epsilon) \sim \Omega^{-1} \sim |\epsilon|^{1/p}.$$

The hyperscaling relation, together with the definition of ν implies

$$G(\epsilon) \sim (r_{\text{corr}})^{-d} \sim |\epsilon|^{d\nu}.$$

Comparing the two expressions for $G(\epsilon)$ we find

$$d\nu = 1/p.$$

From Problem 2c, we have $1/p = 2 - \alpha$; therefore

$$d\nu = 2 - \alpha,$$

the Josephson identity.

[2] Evaporative cooling

(a) The Maxwell distribution is of the form

$$P(v) = \sqrt{\frac{2}{\pi}} \left(\frac{m}{\tau}\right)^{3/2} v^2 e^{-\frac{1}{2}mv^2/\tau}. \quad (2.1)$$

Check the normalization,

$$\begin{aligned} \int_0^\infty dv P(v) &= \sqrt{\frac{2}{\pi}} 2^{3/2} \int_0^\infty dv \left(\frac{m}{2\tau}\right)^{1/2} \frac{1}{2\tau} mv^2 e^{-\frac{1}{2}mv^2/\tau} \\ &= \frac{4}{\sqrt{\pi}} \int_0^\infty dx \left(x^2 e^{-x^2}\right) = 1, \end{aligned} \quad (2.2)$$

where

$$x = \left(\frac{1}{2\tau}mv^2\right). \quad (2.3)$$

Probability that $\frac{1}{2}mv^2 > x_0$ is

$$P(x_0) = \frac{4}{\sqrt{\pi}} \int_{x_0}^\infty dx \left(x^2 e^{-x^2}\right). \quad (2.4)$$

And

$$P(x_0 = 1) = 0.572407, \quad (2.5)$$

which is the fraction of particle that escapes.

- (b) Number of particles with speed in $(v, v + dv)$ is $NP(v)dv$, carrying energy $\frac{1}{2}mv^2NP(v)dv$. Total energy is

$$U = N \int_0^\infty dv P(v) \frac{1}{2}mv^2 = N\tau \frac{4}{\sqrt{\pi}} \int_0^\infty dx x^4 e^{-x^2} = \frac{3}{2}N\tau. \quad (2.6)$$

Energy of particles with $\frac{1}{2\tau}mv^2 > x_0$ is

$$U(x_0) = N\tau \frac{4}{\sqrt{\pi}} \int_{x_0}^\infty dx x^4 e^{-x^2} = \frac{3}{2}N\tau g(x_0), \quad (2.7)$$

where $g(x_0)$ is fraction of energy lost.

$$g(1) = 0.849145 \quad (2.8)$$

- (c) The remaining particles have energy

$$\bar{U} = (1 - g)U = (1 - g)\frac{3}{2}N\tau. \quad (2.9)$$

$$\bar{N} = (1 - f)N \quad (2.10)$$

After evaporation,

$$\bar{U} = (1 - g)\frac{3}{2}N\tau = \frac{3}{2}\bar{N}\bar{\tau} = \frac{3}{2}(1 - f)N\bar{\tau}. \quad (2.11)$$

So,

$$\bar{\tau} = \frac{1 - g}{1 - f}\tau = 0.3528\tau, \quad (2.12)$$

for $x_0 = 1$.

[3] Einstein relation on a lattice

- (a)

$$\frac{\Gamma(s \rightarrow s + 1)}{\Gamma(s + 1 \rightarrow s)} = e^{F\Delta/\tau}. \quad (3.13)$$

So,

$$\begin{aligned} \frac{P_R}{P_L} &= e^{F\Delta/\tau} \rightarrow \\ P_R &= \frac{e^{F\Delta/\tau}}{1 + e^{F\Delta/\tau}}, P_L = \frac{1}{1 + e^{F\Delta/\tau}}. \end{aligned} \quad (3.14)$$

And,

$$P_R - P_L = \frac{e^{F\Delta/\tau} - 1}{e^{F\Delta/\tau} + 1} = \tanh\left(\frac{F\Delta}{2\tau}\right). \quad (3.15)$$

- (b) For $\frac{F\Delta}{\tau} \ll 1$, $P_R - P_L \sim \frac{F\Delta}{2\tau}$. In n steps, the time $t = n\epsilon$, in expectation. The particle hops distance is

$$n\Delta(P_R - P_L) = n\Delta\left(\frac{F\Delta}{2\tau}\right) = n\epsilon\left(\frac{\Delta^2}{2\epsilon}\right)\frac{F}{\tau} = tD\frac{F}{\tau}. \quad (3.16)$$

So the drift speed is

$$\frac{\text{distance}}{\text{time}} = D\frac{F}{\tau}. \quad (3.17)$$

Mobility $b = \frac{v_{drift}}{F} = \frac{F}{\tau}$.

[4] Einstein relation from linear response theory

(a)

$$Z_0 = \sum_a e^{-E_a/\tau} \quad (4.18)$$

and

$$Z_\lambda = \sum_a e^{-(\epsilon_a - \lambda A_a)/\tau} \sim Z_0 + \frac{\lambda}{\tau} Z_0 \langle A \rangle. \quad (4.19)$$

And we know that

$$\sum_a B_a e^{-(\epsilon_a - \lambda A_a)/\tau} \sim Z_0 \langle B \rangle + \frac{\lambda}{\tau} Z_0 \langle AB \rangle. \quad (4.20)$$

Therefore,

$$\langle B \rangle_\lambda = \frac{Z_0 \langle B \rangle + \frac{\lambda}{\tau} Z_0 \langle AB \rangle}{Z_0 + \frac{\lambda}{\tau} Z_0 \langle A \rangle} \sim \langle B \rangle_0 + \frac{\lambda}{\tau} (\langle AB \rangle_0 - \langle A \rangle_0 \langle B \rangle_0). \quad (4.21)$$

(b) Next consider $H \rightarrow H - Fx$. Find drift velocity $\langle v \rangle_F$, where $\langle v \rangle_0 = 0$.

$$\langle v \rangle_F = \frac{F}{\tau} \langle vx \rangle_0 \quad (4.22)$$

[5] The cost of erasure

(a) According to Boltzmann distribution, the probability is proportional to $e^{-E/\tau}$. Thus, we have

$$P(0) = \frac{1}{1 + e^{-\lambda/\tau}} \quad (5.23)$$

$$P(1) = \frac{e^{-\lambda/\tau}}{1 + e^{-\lambda/\tau}} \quad (5.24)$$

(b)

$$dW = P(0)dE_0 + P(1)dE_1 \quad (5.25)$$

$$dW = \frac{e^{-\lambda/\tau}}{1 + e^{-\lambda/\tau}} d\lambda \quad (5.26)$$

(c)

$$W = \int_0^\infty dW = \int_0^\infty \frac{e^{-\lambda/\tau}}{1 + e^{-\lambda/\tau}} d\lambda = \tau \ln 2 \quad (5.27)$$

The result agrees with the Landauer's Principle.